BSM @ LHC

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The giant news of the last year is that we have discovered the Higgs boson! Is it the SM Higgs?



$$\Delta M_H^2 = N_f \frac{\lambda_f^2}{8\pi^2} \left[-\Lambda^2 + 6m_f^2 \log \frac{\Lambda}{m_f} - 2m_f^2 \right]$$

Good reason to expect new physics beyond the Standard Model (SM).

SUSY

Little Higgs

Higgsless Technicolor

Extra dimensions

...others



...others





What was the problem?

Problem I:

Implementing a model was often tedious and error prone.



Terminal – less – 82×38

▲ ▼

OFD Interactions # 2 heavy fermions - 1 light weak gauge boson # FFV (qqZ) dp dp z GZDp QED-HF up up z GZUp QED-HF sp sp z GZDp QED-HF cp cp z GZUp QED-HF bp bp z GZDp QED-HF tp tp z GZTp QED-HF # FFV (llZ) ep- ep- z GZLp QED-HF mup- mup- z GZLp QED-HF tap- tap- z GZLp QED-HF vep vep z GZNp QED-HF vmp vmp z GZNp QED-HF vtp vtp z GZNp QED-HF # FFV (qq'W) - diagonal CKM dp up w- GWFp QED-HF sp cp w- GWFp QED-HF bp tp w- GWTp QED-HF up dp w+ GWFp QED-HF cp sp w+ GWFp QED-HF tp bp w+ GWTp QED-HF # FFV (ll'W) vep ep- w+ GWFp QED-HF vmp mup- w+ GWFp QED-HF vtp tap- w+ GWFp QED-HF ep- vep w- GWFp QED-HF mup- vmp w- GWFp QED-HF tap- vtp w- GWFp QED-HF

```
Terminal — bash — 82×38
                                                                                        V-light
                F-heavy
                            F-heavy
С
      GZDpL =
    - -1d0/2d0*gf(-ee,WMASS,ZMASS,MWP)

    +vZ0f(WMASS,ZMASS,MWP)*vLP0f(WMASS,MWP)**2

    - 1d0/2d0*gtf(-ee,WMASS,ZMASS,MWP)

    +VZ1f(WMASS,ZMASS,MWP)*vLP1f(WMASS,MWP)**2

    - +1d0/6d0*qpf(-ee,WMASS,ZMASS,MWP)

    +vZ2f(WMASS,ZMASS,MWP)

        GZDpR =

    - -1d0/2d0*qtf(-ee,WMASS,ZMASS,MWP)

    +VZ1f(WMASS,ZMASS,MWP)

    - +1d0/6d0*gpf(-ee,WMASS,ZMASS,MWP)

    +vZ2f(WMASS,ZMASS,MWP)

     GZDp(1)=dcmplx(GZDpL,Zero)
      GZDp(2)=dcmplx(GZDpR,Zero)
      write(*,10) 'GZDpL = ',GZDpL
      write(*,10) 'GZDpR = ',GZDpR
      GZUpL =
    - 1d0/2d0*qf(-ee,WMASS,ZMASS,MWP)

    +vZ0f(WMASS,ZMASS,MWP)*vLP0f(WMASS,MWP)**2

    - +1d0/2d0*gtf(-ee,WMASS,ZMASS,MWP)

    +VZ1f(WMASS,ZMASS,MWP)*vLP1f(WMASS,MWP)**2

    - +1d0/6d0*gpf(-ee,WMASS,ZMASS,MWP)

    +vZ2f(WMASS,ZMASS,MWP)

        GZUpR =
    - 1d0/2d0*gtf(-ee,WMASS,ZMASS,MWP)

    +VZ1f(WMASS,ZMASS,MWP)

    - +1d0/6d0*gpf(-ee,WMASS,ZMASS,MWP)

    +vZ2f(WMASS,ZMASS,MWP)

      GZUp(1)=dcmplx(GZUpL,Zero)
      GZUp(2)=dcmplx(GZUpR,Zero)
      write(*,10) 'GZUpL = ',GZUpL
      write(*,10) 'GZUpR = ',GZUpR
```

Problem 2:

Each matrix element generator has its strengths. What if you need more than one? In the past you had to start over.













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Problem 3:

Implementations often did not transfer well to experimentalists.

Problem 3:

Implementations often did not transfer well to experimentalists. It often required modifying the code of the matrix element generator.





MC4BSM 2012 Tutorial

arXiv:1209.0297





	LanHEP	FeynRules	SARAH	
First Released	1996	2008	2008	
Programming Language	С	Mathematica	Mathematica	
General Lagrangian	Yes	Yes	SUSY Only	
Superfields	No	Yes	Yes	
Parameter Running	No	In Progress	Yes	
Aut. Mass Diagonalization	Yes	Yes	Yes	
Spin	0,1/2,1,3/2,2	0,1/2,1,3/2,2 -		
Superfields	_	Chiral, Vector	Chiral, Vector	

	LanHEP FeynRule		SARAH	
CalcHEP	Yes	Yes	Yes	
FeynArts	Yes	Yes	Yes	
MadGraph	In Progress	Yes	Yes	
Sherpa	No	Yes	No	
Whizard	No	Yes	Yes	

Example: IDP

CPC 184(2013) 1729-1769

$$H_1 = \begin{pmatrix} 0 \\ \langle v \rangle + h/\sqrt{2} \end{pmatrix} , \quad H_2 = \begin{pmatrix} \widetilde{H}^+ \\ (\widetilde{X} + i\widetilde{H}_3)/\sqrt{2} \end{pmatrix}$$

$$\mathcal{L} = \mathcal{L}_{SM} + D^{\mu} H_2^* D_{\mu} H_2 - \mu_2^2 |H_2^2|^2 -\lambda_2 |H_2|^4 - \lambda_3 |H_1|^2 |H_2|^2 - \lambda_4 |H_1^{\dagger} H_2|^2 - \lambda_5 Re[(H_1^{\dagger} H_2)^2]$$

```
parameter MHX=111,MH3=222,MHC=333. % Declaration of new masses
parameter laL=0.01, la2=0.01. % Declaration of new couplings
%mu^2 as a function of masses
parameter mu2=MHX**2-laL*(2*MW/EE*SW)**2.
% constraints for couplings
parameter la3=2*(MHC**2-mu2)/(2*MW/EE*SW)**2.
parameter la5=(MHX**2-MH3**2)/(2*MW/EE*SW)**2.
parameter la4=2*laL-la3-la5.
```

FeynRules

```
scalar '~H3'/'~H3':('odd Higgs',pdg 36, mass MH3, width wH3 = auto).
scalar '~H+'/'~H-':('Charged Higgs',pdg 37,mass MHC,width wHC=auto).
scalar '~X'/'~X':('second Higgs',pdg 35,mass MHX,width wHX=auto).
let h2 = { -i*'~H+', ('~X'+i*'~H3')/Sqrt2 },
    H2 = { i*'<sup>H-'</sup>, ('<sup>X</sup>'-i*'<sup>H3'</sup>)/Sqrt2 }.
                                     FeynRules
M$ClassesDescription = {
                                               S[23] == {
S[21] == {
                                                   ClassName -> HC,
   ClassName -> X,
                                                   SelfConjugate -> False,
   SelfConjugate -> True,
                                                   Mass -> {MHC,333},
                                                   Width \rightarrow {wHC,0},
                  -> \{MHX, 111\},
   Mass
                   -> \{wHX, 0\},\
                                                   QuantumNumbers \rightarrow \{Q \rightarrow 1\},\
   Width
   PDG
                  -> 35,
                                                   PDG
                                                           -> 37,
                                                   ParticleName -> "~H+",
   ParticleName -> "~X",
                                                   AntiParticleName -> "~H-",
                   -> "second Higgs"
   FullName
                                                   FullName -> "Charged Higgs"
  },
                                                 },
S[22] == {
                                               S[24] == {
                   -> H3,
   ClassName
                                                   ClassName -> h2,
    SelfConjugate -> True,
                                                   Unphysical -> True,
                  -> {MH3,222},
    Mass
                                                   Indices -> {Index[SU2D]},
                   -> {wH3,0},
    Width
                                                   FlavorIndex -> SU2D,
                   -> 36,
    PDG
                                                   SelfConjugate -> False,
                   -> "~H3",
    ParticleName
                                                   QuantumNumbers \rightarrow {Y \rightarrow 1/2},
                   -> "odd Higgs"
    FullName
                                                   Definitions \rightarrow { h2[1] \rightarrow -I HC,
  },
                                                                h2[2] \rightarrow (X + I H3)/Sart[2]
```

FeynRules

```
LIDM1 = DC[h2bar[ii], mu] DC[h2[ii], mu];
LIDM2 = -mu2^2 h2bar[ii] h2[ii];
LIDM3 = -la2 h2bar[ii] h2[ii] h2bar[jj] h2[jj];
LIDM4 = -la3 Phibar[ii] Phi[ii] h2bar[jj] h2[jj];
LIDM5 = -la4 h2bar[ii] Phi[ii] Phibar[jj] h2[jj];
LIDM6 = -la5/2 h2bar[ii] Phi[ii] h2bar[jj] Phi[jj];
LIDM7 = HC[LIDM6];
LIDM = LIDM1 + LIDM2 + LIDM3 + LIDM4 + LIDM5 + LIDM6 + LIDM7;
```

lhep <source file> -ca -evl 2

FeynRules

\$FeynRulesPath = "<FR path>";
SetDirectory[\$FeynRulesPath];
<< FeynRules';</pre>

SetDirectory[<IDM path>];
LoadModel["SM.fr", "IDM.fr"];

WriteCHOutput[LSM, LIDM]

Automatic Mass Diagonalization

Eur.Phys.J. C73 (2013) 2325

M\$MixingsDescription = {
 Mix["l1"] == { options1 },
 Mix["l2"] == { options2 },

Automatic Mass Diagonalization

Eur.Phys.J. C73 (2013) 2325

Automatic Mass Diagonalization

Eur.Phys.J. C73 (2013) 2325

Mix["AZmix"] == {
 MassBasis -> {A, Z},
 GaugeBasis -> {B, Wi[3]},
 MixingMatrix -> UW,
 BlockName -> WEAKMIX

MEG Hack Not Required

These model implementations can be used just like built in models. They do not require modification of the matrix element generator code!

Validation

Validation

- Check Hermiticity.
- Check Feynman rules with literature .
- Check gauge invariance.
- Check consistency between supported matrix element generators.
- Check distributions.





Remove

Remove

Remove

Remove

Remove

-

New Standard Model

Standard Model (current)

Standard Model (development)

New Standard Model v4

Standard Model

DICD



Create New MEG Model Files



The probability of this being a statistical fluctuation is 0.0%.

	\sqrt{s}	p _{Tcut}	Best	CH(u)	WO2(u)	X ²
$\sim n1$, $\sim n1 \rightarrow s$, S	3202.0	800.5	7.104E-04	7.103E-04	7.106E-04	2.0E+00
$m \hspace{0.2cm}, \hspace{0.2cm} M \hspace{0.2cm} \rightarrow \hspace{0.2cm} s \hspace{0.2cm}, \hspace{0.2cm} S$	400.0	100.0	5.070E-01	5.071E-01	5.069E-01	1.5E+00
$t \hspace{0.1in}, \hspace{0.1in} T \hspace{0.1in} \rightarrow \hspace{0.1in} s \hspace{0.1in}, \hspace{0.1in} S$	2754.0	688.5	2.444E-01	2.444E-01	2.442E-01	1.5E+00
$u \hspace{0.2cm}, \hspace{0.2cm} U \hspace{0.2cm} \rightarrow \hspace{0.2cm} d \hspace{0.2cm}, \hspace{0.2cm} S$	400.0	100.0	5.322E-01	5.322E-01	5.325E-01	1.2E+00
${\sim}n2$, ${\sim}n2$ \rightarrow ${\sim}n2$, ${\sim}n2$	3200.0	800.0	1.225E-02	1.225E-02	1.224E-02	1.1E+00
t , T \rightarrow d , D	2752.0	688.0	2.455E-01	2.455E-01	2.453E-01	1.0E+00
t , $T \rightarrow Z$, H1	2221.0	555.25	3.203E-01	3.203E-01	3.207E-01	9.1E-01

Model Databases
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•		\bigcirc	feynrules.irmp.ucl.ac.be/wiki/ModelDatabaseMainPage	Reader 🖒 Search	
×			ModelDatabaseMainPage – FeynRules		+

FeynRules model database

This page contains a collection of models that are already implemented in FeynRules. For each model, a complete model-file is available, containing all the information that is needed, as well as the Lagrangian, as well as the references to the papers were this Lagrangian was taken from. All model-files can be freely downloaded and changed, serving like this as the starting point for building new models. A TeX-file for each model containing a summary of the Feynman Rules produced by FeynRules is also available.

The Standard model model-file is already included in the distribution of the FeynRules, but it can also be downloaded independently from the corresponding link below.

We encourage model builders writing a FeynRules implementation of their model to make their model file(s) public in the FeynRules model database, in order to make them useful to a community as wide as possible. For further information on how to make your model implementation public via the FeynRules model database, please send an email to

- duhrc@...
- neil@...
- fuks@...
- cdegrand@...

Available models

Standard Model	The SM implementation of FeynRules, included into the distribution of the FeynRules package.
Simple extensions of the SM (10)	Several models based on the SM that include one or more additional particles, like a 4th generation, a second Higgs doublet or additional colored scalars.
Supersymmetric Models (4)	Various supersymmetric extensions of the SM, including the MSSM, the NMSSM and many more.
Extra-dimensional Models (4)	Extensions of the SM including KK excitations of the SM particles.
Strongly coupled and effective field theories (5)	Including Technicolor, Little Higgs, as well as SM higher-dimensional operators.
Miscellaneous (0)	

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Sear	ch in HEPMDB	٩,	Show All Models					
Sea	rch Models :: Results for	r [AII]						
1.	E6SSM-12.02 [2012-11-08 13 J.P Hall, P. Svantesson Electroweak scale CalcHEP mode MSSM and USSM. In addition to	5:23:02] hepmdb:11 el of the E6SSM [v12. the MSSM particle co	12.0106 02]. The E6SSM (King. et al. 2006) is intent, this version of the E6SSM cont	an extension of	the			
2.	ESM [2012-10-13 07:35:48] h M.V. Chizhov	1epmdb:1012.0103						
3.	The model is described in arXiv: Next to Minimal Walking Tec hepmdb:1012.0102	1005.4287 :hnicolour Model (N	MWTC) in CalcHEP format [2012-	10-10 08:56:07]				
	The SU(3) theory with two flavor Walking Technicolor (NMWT) in t	rs in the two-index sy the CalcHEP format, s	matti J arvinen, and Francesco Sann. mmetric representation which the see 1. "Technicolor Walks at the LHC.	Next to Minimal	IKNOV			
4.	LQ 3rd generation for CalcHI A.Belyaev, A.Pukhov	EP [2012-06-20 13:	59:30] hepmdb:0612.0078					
5.	 LQ 2nd generation for CalcH	EP [2012-06-20 13:	58:34] hepmdb:0612.0077					

A.Belyaev, A.Pukhov

Future



$$\log(1.27) = 0.27 - \frac{1}{2}0.27^2 + \frac{1}{3}0.27^3 - \frac{1}{4}0.27^4 + \cdots$$

$$\log(1.27) = 0.27 - \frac{1}{2}0.27^2 + \frac{1}{3}0.27^3 - \frac{1}{4}0.27^4 + \cdots$$

$$\cos(0.33) = 1 - \frac{1}{2!} 0.33^2 + \frac{1}{4!} 0.33^4 + \cdots$$

 $\frac{1}{0.57} = 1 + 0.43 + 0.43^2 + 0.43^3 + 0.43^4 + \cdots$

•

RPN Calalator
0.0000
Image of mass loss in any set of the set of th

$$\log(1.27) = 0.1038\cdots$$

$$\cos(0.33) = 0.9460 \cdots$$

$$\frac{1}{0.57} = 1.754\cdots$$

Why use technology?

- Can you calculate it on your own? Yes!!!
 - Should you learn how to do it? Yes.
- During research, why waste your time on the algorithms?
 - We could do so much more physics if we let computers do the algorithms!
- Rather concentrate on the new physics!
- Humans are good at creativity!
- Computers are good at algorithms!
 - Let the computers do the algorithms once they are mature!

We use technology to Do more!



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Insert vevs





Insert vevsExpand Lagrangian



Insert vevs

- •Expand Lagrangian
- •Collect quadratic terms



- Insert vevs
- Expand Lagrangian
- Collect quadratic terms
- Diagonalize mass matrices



- Insert vevs
- •Expand Lagrangian
- •Collect quadratic terms
- Diagonalize mass matrices
- Rotate fields to mass basis



Insert vevs

- •Expand Lagrangian
- Collect quadratic terms
- Diagonalize mass matrices
- Rotate fields to mass basis
- Calculate Feynman diagrams



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- •Implement Feynman rules into CH, FA, MG, SH, WO



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- Implement Lagrangian into FR, LH, SARAH



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- •Expand Lagrangian
- Collect quadratic terms
- Diagonalize mass matrices
- Rotate fields to mass basis
- •Calculate Feynman diagrams
- •Implement Feynman rules into CH, FA, MG, SH, WO
- Implement Lagrangian into FR, LH, SARAH
- Do calculations



- Insert vevs Algorithmic!
- Expand Lagrangian Algorithmic!
- Collect quadratic terms Algorithmic!
- Diagonalize mass matrices Algorithmic!
- •Rotate fields to mass basis Algorithmic!
- Calculate Feynman diagrams Algorithmic!
- Implement Feynman rules into CH, FA, MG, SHgoMOmic!
- •Implement Lagrangian into FR, LH, SARAH Algorithmic!
- Do calculations Some algorithmic!

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He worked very hard to improve the telescope and ...





He worked very hard to improve the telescope and then he pointed it at the skies and discovered craters on the moon, moons orbiting jupiter, and many others!





He worked very hard to improve the telescope and then he pointed it at the skies and discovered craters on the moon, moons orbiting jupiter, and many others!

We have the same opportunity in our day!

Introducing



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U	-Lagrangian $-\frac{1}{4}G_{\mu\nu}G^{\mu\nu}+6$	$ \theta_0 G_{\mu\nu} \widetilde{G}^{\mu\nu} + -\frac{1}{4} W_{\mu\nu} W^{\mu} $	$\nu + \theta_1$	$W_{\mu\nu}\widetilde{W}$	$\mu\nu + -\frac{1}{4}$	Β _{μν} Β ^μ	$\nu + i \overline{Q_L}$	γ _μ D ^μ Q ₁	$L + i \overline{u_R} \gamma_{\mu} D$	$^{\mu}u_{R}+$		1
1	$i \overline{d_R} \gamma_{\mu} D^{\mu} d_R +$	$i\overline{L_L}\gamma_{\mu}D^{\mu}L_L+i\overline{e_R}\gamma_{\mu}D^{\mu}$	e_R +1	$D_{\mu} \Phi D^{\mu}$	Φ^*	-		-				
>_	$\mu_{r2} \Psi \Psi$ $\lambda_{r3} \Phi \Phi \Phi^* \Phi^* +$											
	$\lambda_{r4} \Phi \Phi \Phi^* \Phi^*$		_	_								
· •	$y_5 \Phi \overline{u_R} Q_L + y_5^*$	$\Phi^* \overline{Q_L} u_R + y_6 \Phi^* d_R Q_L + y_6 \Phi^* d_R$	${}_{6}^{*}\Phi Q$	$_L d_R + y_7$	$\Phi^* \overline{e_R} L_L$	+						
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1	$i\frac{4}{d_R}\gamma_{\mu}D^{\mu}d_R+i\overline{L_L}\gamma_{\mu}D^{\mu}L_L+i\overline{e_R}\gamma_{\mu}D$	$^{\mu}e_{R}+$	$D_{\mu}\Phi D^{\mu}$	Φ^*	μυ		μ- ι				
	$\mu_{r2} \Phi \Phi^*$										
<u>></u>	$\lambda_{r3} \Phi \Phi \Phi^* \Phi^* +$										
10	$y_{5}\Phi \overline{u_{R}}Q_{L} + y_{5}^{*}\Phi^{*}\overline{Q_{L}}u_{R} + y_{6}\Phi^{*}\overline{d_{R}}Q_{L} +$	$y_6^* \Phi \overline{Q}$	$\overline{D_L} d_R + y_7$	$\Phi^* \overline{e_R} L_L$	$+y_{7}^{*}\Phi$	$\overline{L_L}e_R$ +	$V_8 \Phi \Phi \overline{L_L^c}$	$L_L + y_8^* \Phi^*$	$\Phi^* \overline{L_L} L_L^c +$		
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	$\frac{4}{i\overline{d_R}} v_{\mu} D^{\mu} d_R + i\overline{L_I} v_{\mu} D^{\mu} L_I + i\overline{e_R}$	$v_{\mu}D^{\mu}e_{\mu}$	$+D \Phi L$	$\Phi^{\mu} \Phi^{*} + 1$	4 - μν - D H ₂ L	$D^{\mu}H$	L*	$\mu = - q_L \cdots = q_R \mu = - q_R \cdot$		
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12	$\int d^2\theta d^2\overline{\theta} H^{\dagger} \exp[q_1W + q_2B]H_1 + -\frac{1}{2}\int d^2\theta d^2\overline{\theta} H^{\dagger} \exp[q_1W + q_2B]H_1 + -\frac{1}{2}\int d^2\theta d^2\overline{\theta} H^{\dagger} \exp[q_1W + q_2B]H_2 + -\frac{1}{2}\int d^2\theta d^2\overline{\theta} \exp[q_1W + q_2B]H_2 + -\frac{1}{2}\int d^2\theta \exp[q_1W + q_2B]H_2 + -1$	$d^2 \theta G G = \frac{1}{2}$	$\int d^2 \overline{\theta} G$	$\int \mathbf{C}^{\dagger} \mathbf{C}^{\dagger} \mathbf{+} \mathbf{-}$	1 d ²	WW =	$-\frac{1}{d^2 \overline{\Theta} W^{\dagger}}$	W^{\dagger} +		
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0	$y_5 \int d^2 \overline{\theta} d_c^{\dagger} d_c^{\dagger} u_c^{\dagger} + y_5^* \int d^2 \theta u_c d_c d_c + y_6 \int d^2 \theta u_c d_c d_c d_c + y_6 \int d^2 \theta u_c d_c d_c d_c + y_6 \int d^2 \theta u_c d_c d_c d_c d_c + y_6 \int d^2 \theta u_c d_c d_c d_c d_c d_c d_c d_c d_c d_c d$	$d^2\overline{\theta}e_c^{\dagger}L^{\dagger}l$	$L^{\dagger} + y_6^* \int$	d ² θLL	$e_c + y_7$	d²∂H	$I_d^{\dagger} e_c^{\dagger} L^{\dagger} + y_7^*$	$\int d^2\theta L e_c H_d + y_8$	$d^2\overline{\theta}H^{\dagger}_dH^{\dagger}_dH^{\dagger}_de^{\dagger}_c$	
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Galileo : Current

- Supports any semisimple compact Lie algebra (symmetry).
- Supports fields of spin 0,1/2,1.
- Supports superfields.
- Automatically generates the Lagrangian.
- Core library + GUI wrapper.

Galileo : Plans

- Core and GUI need more polishing.
- Expand Lagrangian.
- Symmetry breaking.
- Mass matrix diagonalization and rotation to physical basis.
- Save/read.
- Export to FeynRules (for further analysis).



1997: Deep blue beats the current world champion, Garry Kasparov at chess!



In 2011, Watson beat the world champions at Jeopardy!

In 2025, ...

Charles Xavier Thomas de Colmar invented the first commercially successful mechanical calculator in 1820. It was 100 years before mechanical calculators gave way, in the 1930s, to electromechanical calculators, which then quickly gave way to the first general-purpose electronic computer, ENIAC, in 1946. By 1965, Gordon Moore was predicting that engineers would be able to double the number of components on a microchip every two years (and by 1968, he co-founded Intel to help them do so).

Just as Moore predicted, computers continue to become exponentially faster, while their components have become much cheaper. William Nordhaus, an economist at Yale University, examined hundreds of devices—from the first computer to the Apple II to modern PCs—and determined how many basic calculations they could perform every second.



Let's be part of the next revolution in science.

Summary

- Discovery of a Higgs-like state!
- We still expect to find more new physics BSM.
- It is much easier/safer to implement BSM models into matrix element generators.
 - Several choices for implementation (LanHEP, FeynRules, SARAH).
 - Export to many MEGs (CalcHEP, FeynArts, MadGraph, Sherpa, Whizard).
 - No need to modify MEG code.
 - Improved validation available.
- Model databases available (FeynRules and HEPMDB).
- Future: Galileo will make the situation even better for the study of BSMs.